

Stability of low-spin phase D in the lowermost mantle

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The Earth's mantle constitutes half of the planet's mass and may contain several times the amount of water as the combined surface reservoirs, dispersed within high-pressure silicates and melts. Water is cycled between the crust and mantle *via* plate tectonics, but there is also evidence for primordial reservoirs in the deeper mantle (Hallis *et al.*, 2016). Evidence for H₂O-bearing phases trapped as inclusions in diamond emanating from the transition zone (410–660 km depth), and deeper still from the top of the lower mantle (~700–800 km depth), suggest that hydrous melts and minerals are present throughout most of the Earth's mantle (Pearson *et al.*, 2014; Palot *et al.*, 2016). A major question is: what are the dominant host minerals for water in the lower mantle, *i.e.* at depths of around 660–2900 km? Among all of the known dense hydrous magnesium silicates, phase D (MgSi₂O₄(OH)₂) and phase H (MgSiO₄H₂) are expected to have the highest stability in the lowermost mantle. Ohtani *et al.* (2014) found phase H stable in the peridotite system from 35–65 GPa at elevated temperatures, with a possibly expanded stability upon dissolution of the isostructural δ -AlOOH component. The magnesium end-member of phase D breaks down at 20–30 GPa at temperatures above ~1300–1400 °C, and at around 50 GPa below 1000 °C (Frost and Fei, 1998). However, a spin transition was recently observed in iron bearing phase D at around 40–65 GPa (Chang *et al.*, 2013), and the stability of the low spin phase may be stable to much lower depths. We are investigating the high *P–T* stability of Fe- and Al-bearing phase D at deeper mantle conditions using the laser-heating setup at GSECARS. Our results place new constraints on the stability of this family of hydrous minerals, with implications for which hosts carry water into the deep Earth.